

Onset of instability for the Alber equation, and applications for rogue waves

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Recent progress on rogue waves and their applications

This talk



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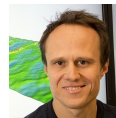
*Mariya
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*Gerassimos
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*Themistoklis
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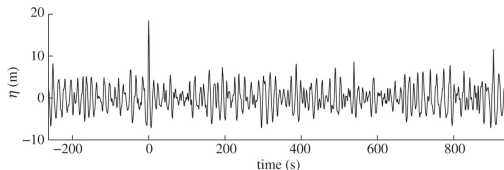
*Odin
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- A. Athanassoulis, G. Athanassoulis and T. Sapsis, **Journal of Ocean Engineering and Marine Energy** (2017)
- A. Athanassoulis, G. Athanassoulis, M. Ptashnyk and T. Sapsis, **Kinetic and Related Models** (2020)
- A. Athanassoulis and O. Gramstad, **Fluids** (2021)
- A. Athanassoulis and I. Kyza, **Water Waves** (2023)
- A. Athanassoulis, I. Kyza and F. Karakatsani, to appear in **Applied Numerical Mathematics** (arXiv:2506.06879)
- A. Athanassoulis, **arXiv:2604.16998**

Two main meanings of “Rogue waves” in this session

(A) Oceanography

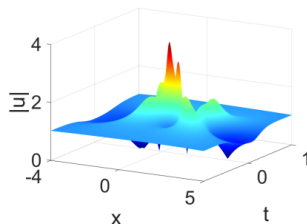
- A **localized event** far larger than the RMS size in a stochastic wave field.
- Observed in nature, in wavefields that are characterized through their **power spectrum (not explicitly known)**.
- **Modulation instability** is relevant.



Draupner platform, the North Sea

(B) Integrable equations

- A solution of a nonlinear dispersive equation with **sharply localized features**, on a background.
- The background is **explicitly known** (plane wave; more recently elliptic / periodic etc).
- **Modulation instability** is relevant.



Sasa-Satsuma RW [Wu et al.]

Deriving the Alber equation (sketch)

1. **Envelope.** A envelope for unidirectional sea state $\rightarrow$ focusing cubic NLS:

$$i\partial_t A + p \Delta A + \frac{q}{2}|A|^2 A = 0.$$

2. **Stochastic modeling: Second moments.** Form two-space autocorrelation $R(x, y, t) = \mathbb{E}[A(x, t)\overline{A(y, t)}]$. A Gaussian (BBGKY) closure gives, to leading order,

$$i\partial_t R + p(\Delta_x - \Delta_y)R + q(R(x, x, t) - R(y, y, t))R = 0. \quad [\text{Alber 1978}]$$

3. **Constitutive assumption: quasi-homogeneity.** Seas are nearly homogeneous, so split

$$R(x, y, t) = \underbrace{\Gamma(x - y)}_{\text{known background}} + \underbrace{u(x, y, t)}_{\text{inhomogeneity}}$$

$$i\partial_t u + p(\Delta_x - \Delta_y)u + q(V(x, t) - V(y, t))(\Gamma(x - y) + u) = 0$$
$$V(x, t) = u(x, x, t)$$

- Γ (equivalently the power spectrum $P = \mathcal{F}\Gamma$) is **data measured from the field**.
- Equation structure related to a *von Neumann equation with δ self-interaction* / mixed-state NLS.

The difficulties

The Alber equation

$$i\partial_t u + p(\Delta_x - \Delta_y)u + q(u(x, x, t) - u(y, y, t))(\Gamma(x - y) + u) = 0$$

is a dispersive nonlinear equation. But:

- Hyperbolic Laplacian not sign-definite $\Rightarrow$ **lose energy bounds**
- trace operator $u(x, y, t) \mapsto V(x, t) = u(x, x, t) \Rightarrow$ **lose $\frac{d}{2}$ derivatives**
- $\partial_t u(x, y, t)$ depends on $u(x, y, t)$, $u(x, x, t)$, $u(y, y, t) \Rightarrow$ **non-locality**

The bifurcation

The Alber equation can be re-written as

$$i\partial_t u + p(\Delta_x - \Delta_y)u + q(u(x, x, t) - u(y, y, t))\Gamma(x - y) = q(u(y, y, t) - u(x, x, t))u$$
$$i\partial_t u + \mathcal{L}_\Gamma u = \mathcal{Q}(u, u)$$

- The linear problem undergoes a bifurcation depending solely on Γ [Alber1978; A. et al. 2020]. $\Rightarrow$ **MI may be ON or OFF!**
- Does fully nonlinear stability follow the linear stability analysis?
No spectral gap $\Rightarrow$ not automatic
- What happens in the fully nonlinear problem in the unstable regime?
 $\Rightarrow$ **ramifications for Oceanic Rogue Waves**

Global well-posedness

Operator-valued (mixed-state) form:

$$i\partial_t \gamma + p[\Delta, \gamma] + q[V_{\rho_\gamma}, \gamma] = 0, \quad \rho_\gamma(x, t) = K_\gamma(x, x, t),$$

with $\gamma = \gamma^* \geq 0$ trace-class, kernel $K_\gamma(x, y, t) = \Gamma(x - y) + u(x, y, t)$, and V_f the operator of multiplication by f .

Theorem (GWP in $H^1\mathfrak{S}^1(\mathbb{T})$; both signs, no smallness [A. arXiv:2604.16998])

Let $pq \neq 0$ and $\gamma_0 \in H^1\mathfrak{S}^1(\mathbb{T})$ be self-adjoint and **non-negative**. Then the Alber equation has a unique global solution $\gamma \in C(\mathbb{R}; H^1\mathfrak{S}^1(\mathbb{T}))$, conserving mass $\|\gamma\|_{\mathfrak{S}^1}$, $\|\gamma\|_{\mathfrak{S}^2}$ and energy $E(\gamma)$, with a uniform bound $\|\gamma(t)\|_{H^1\mathfrak{S}^1} \leq \bar{Y}$ for all t . Higher regularity $H^s\mathfrak{S}^1$ ($s \geq 2$) propagates.

Ingredients (high level):

- Schatten–Sobolev spaces $H^s\mathfrak{S}^1$: control the position density ρ_γ **without losing derivatives**.
- **Energy conservation** + Hoffmann–Ostenhof + Gagliardo–Nirenberg:

$$\|\rho_\gamma\|_{L^2}^2 \leq C \|\gamma\|_{\mathfrak{S}^1}^{3/2} \sqrt{\text{tr}(-\Delta\gamma(t))} + C \|\gamma\|_{\mathfrak{S}^1}^2 \implies \text{tr}(-\Delta\gamma(t)) \leq A + B \sqrt{\text{tr}(-\Delta\gamma(t))}$$

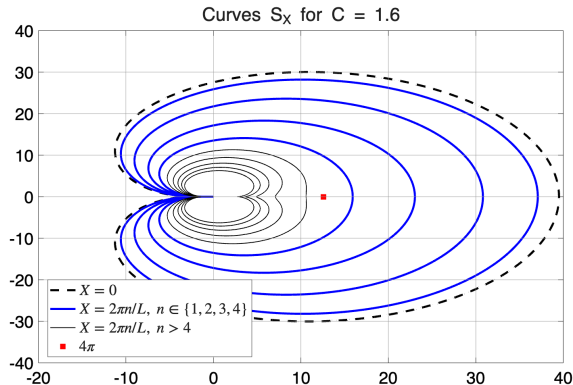
- Non-negativity ($\gamma \geq 0$) is a *structural* requirement (roughly reflecting the non-negativity of the background power-spectrum $P = \mathcal{F}\Gamma$).

Linear stability: the Penrose condition

The wavenumber X_* is unstable iff

$$\exists \omega_*, \operatorname{Re} \omega_* > 0 : \quad \mathbb{H} \left[\frac{P(k + \frac{X_*}{2}) - P(k - \frac{X_*}{2})}{X_*} \right] (\omega_*) = \frac{4\pi p}{q},$$

where $\mathbb{H}$ is the Hilbert transform (acting in k) and $P = \mathcal{F}\Gamma$ the power spectrum.



Instability is read off the background spectrum $P = \mathcal{F}\Gamma$:

$$S_X(t) = \mathbb{H}[D_X P](t) - i D_X P(t)$$

- Wavenumber X is **unstable** $\iff S_X$ winds around $4\pi p/q$.
- *Baseband*: only $|X| < X_{\max}$ unstable.
- $2\pi/X_{\max} \rightarrow$ lengthscale of the coherent structure
- $\max \operatorname{Re} \omega_* \rightarrow$ rate of growth.

Why a new numerical method?

- Despite its standing in oceanography, the Alber equation had **one** bespoke numerical method (Stiassnie–Regev–Agnon 2008; Ribal et al. 2013):
forward Euler + 2nd-order FD + a non-standard far-field boundary condition fitted to each spectrum.
- We aim for structure-preserving; 2^{nd} -order in time + high order in space; easy to implement

Idea

- Frame the Alber equation as an NLS variant on $\mathbb{T}$

$$i\partial_t u + p(\Delta_x - \Delta_y)u + q(u(x, x, t) - u(y, y, t))(\Gamma(x - y) + u) = 0$$

- Use a spatial discretization that allows “silent trace” (spectral / FD on symmetric grid)
- Devise a Besse-type scheme (relaxation Crank-Nicolson) [Besse 2004; Besse et al. 2021]
 $\Rightarrow 2^{nd}$ **order in time** $\Rightarrow$ **linearly implicit** $\Rightarrow$ **structure preserving**

The scheme [A., Kyza & Karakatsani, APNUM 2026]

Auxiliary variable $\phi = u(x, x) - u(y, y)$ on a *staggered* time grid:

$$\frac{1}{2} \left(\Phi^{n+\frac{1}{2}} + \Phi^{n-\frac{1}{2}} \right) = U^n(x, x) - U^n(y, y),$$

$$i \frac{U^{n+1} - U^n}{\tau} + p(\Delta_x - \Delta_y) U^{n+\frac{1}{2}} + q \Phi^{n+\frac{1}{2}} (\Gamma + U^{n+\frac{1}{2}}) = 0.$$

- 2nd order in time, **linearly implicit** (explicit in the nonlinearity).
- 4th-order finite differences in space.
- **Structure-preserving**: discrete L^2 balance law
- Two invariants conserved at the discrete level.

Balance Law carries over to the discrete level:

$$\text{Continuous } \|u\|_{L^2} \text{ Balance Law } \frac{d}{dt} \|u\|_{L^2}^2 = 2 \operatorname{Re} [iq \langle (V(x) - V(y)) \Gamma(x-y), u \rangle]$$

$$\text{Discrete } \|U^n\|_{L^2} \text{ Balance Law } \frac{\|U^{n+1}\|_{L^2}^2 - \|U^n\|_{L^2}^2}{\tau} = 2 \operatorname{Re} [iq \langle \Phi^{n+\frac{1}{2}} \Gamma(x-y), U^{n+\frac{1}{2}} \rangle]$$

Implementation

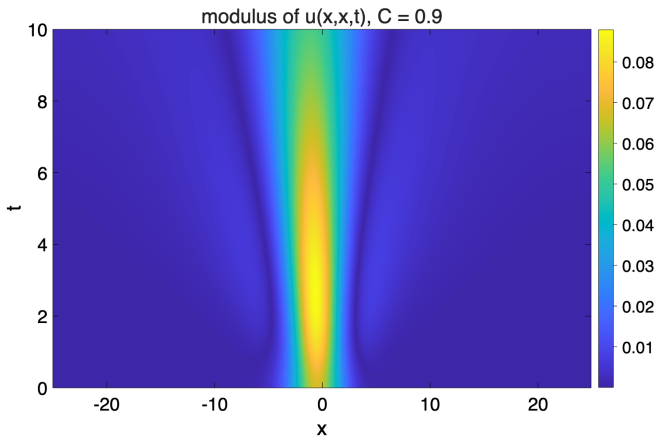
- Uniform $N \times N$ grid on $[-\frac{L}{2}, \frac{L}{2}]^2$, mesh size $h = L/N$
- arrays $U_{i,j}^n$ approximating the solution at (x_i, y_j, t^n) and $\Phi_{i,j}^{n+\frac{1}{2}}$ at the nodes $(x_i, y_j, t^{n+\frac{1}{2}})$.
- $\text{vec}(\cdot)$ is canonical vectorization of an $N \times N$ array into a length- N^2 vector
- $\mathfrak{D}_H$: sparse $N^2 \times N^2$ matrix for the hyperbolic Laplacian with periodic BCs
- $\mathfrak{F}^{n+\frac{1}{2}} = \text{diag}(\text{vec}(\Phi_{i,j}^{n+\frac{1}{2}}))$: pointwise product by $\Phi^{n+\frac{1}{2}}$; I : identity.

Each step: update Φ , solve one sparse linear system for $U^{n+\frac{1}{2}}$, extract U^{n+1} :

$$\begin{aligned}\Phi_{i,j}^{n+\frac{1}{2}} &= 2(U_{i,i}^n - U_{j,j}^n) - \Phi_{i,j}^{n-\frac{1}{2}}, \\ \left(I - i\frac{p\tau}{2}\mathfrak{D}_H - i\frac{q\tau}{2}\mathfrak{F}^{n+\frac{1}{2}}\right) \text{vec}(U_{i,j}^{n+\frac{1}{2}}) &= \text{vec}(U_{i,j}^n + i\frac{q\tau}{2}\Gamma(x_i - y_j)\Phi_{i,j}^{n+\frac{1}{2}}), \\ U_{i,j}^{n+1} &= 2U_{i,j}^{n+\frac{1}{2}} - U_{i,j}^n.\end{aligned}$$

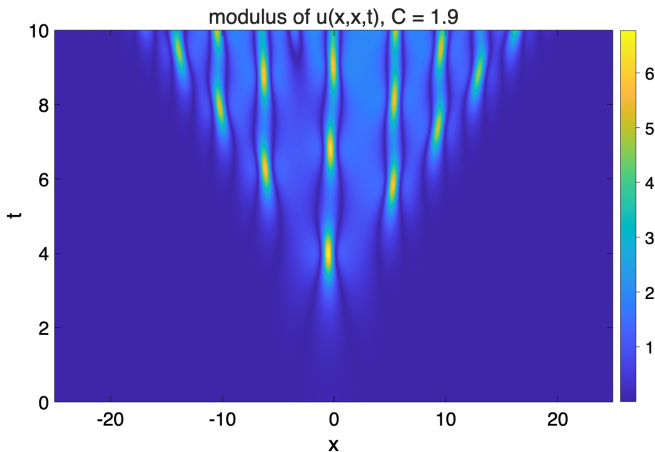
Structure: $M = I - i\frac{\tau}{2}(p\mathfrak{D}_H + q\mathfrak{F}^{n+\frac{1}{2}}) \Rightarrow M$ normal, invertible $\forall \tau$, $\|M^{-1}\|_2 \leq 1$.

Stable regime: nonlinear stability



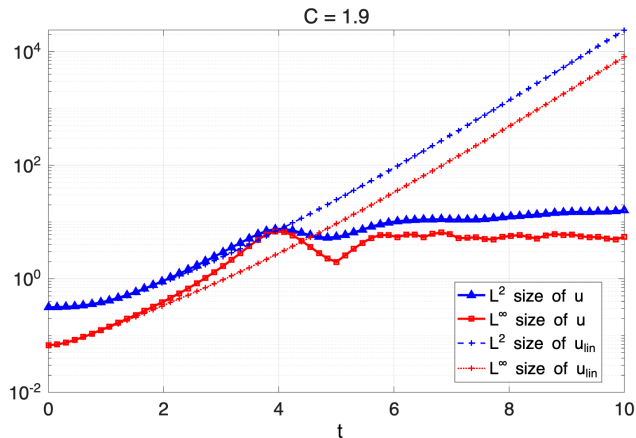
- $|u(x, x, t)|$ for a Penrose-stable spectrum
- No meaningful growth in time – the inhomogeneity simply **disperses**.
- The background Γ would naively act as a source, yet u stays small.
- Proving smallness for all time on $\mathbb{T}$ is open.

Unstable regime: the space–time lattice



- $|u(x, x, t)|$ for a (strongly) Penrose-unstable spectrum
- Stage 1: exponential growth to a single localised coherent structure, $O(1)$, peaking ≈ 6 (initially $O(10^{-2})$).
- Stage 2: **no further L^∞ growth**; a lattice of similar structures spreads outward in a space–time cone.
- *Attracting pattern*: eyeball-robust to the initial u_0 , depends on p, q, Γ .
- Long-time limit open; on $\mathbb{T}$ expected to be different than $\mathbb{R}$.

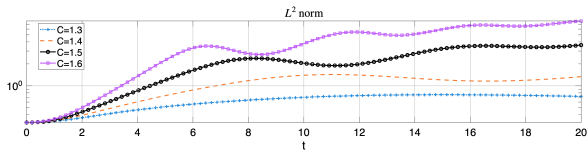
Linear theory: right about what it can see



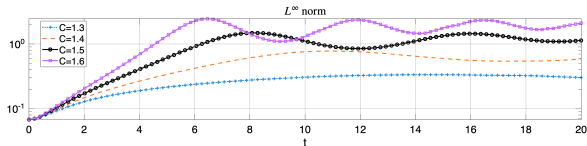
Norms of the full (u) vs. linearised (u_{lin}) solution for a Penrose-unstable problem.

- Linear stability analysis nails the **onset** and the **initial growth rate**.
 - It completely **misses** the maximum amplitude, the coherent structure, and the relevant timescales.
- ⇒ The nonlinear problem must be solved to say anything about extremes.

The bifurcation is gradual – not a switch at onset



L^2 growth vs. time, several C .



L^∞ growth vs. time, several C .

- $C \approx 1$ is the critical point for the bifurcation.
- Linear instability is present for $C > 1$, but for *intermediate* intensities the maxima **plateau** at values negligible next to the original background.
- **Large growth of u compared to $u_0 \neq$ events that dominate Γ**
- For a physically relevant extreme event to actually form, *sufficiently strong* instability is required.

Two amplification factors – and why the distinction matters

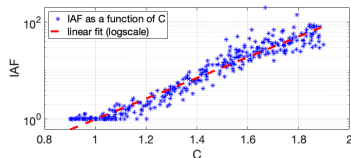
$$\text{IAF} = \frac{\max_{t,x,y} |u|}{\max_{x,y} |u_0|}$$

- **IAF**: how much the inhomogeneity grows (instability strength; numerical difficulty).
- Large IAF is **necessary but not sufficient** for large TAF.

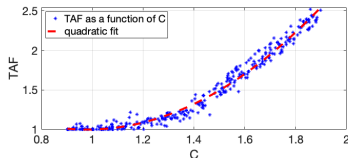
$$\text{TAF} = \frac{\max_{t,x,y} |u + \Gamma|}{\Gamma(0)}$$

- **TAF**: how large the sea state gets *relative to the background* – the physical quantity.
- Oceanic Rogue Wave guaranteed for $\text{TAF} > 2.84$ [Ribal et al., JFM, 2013]

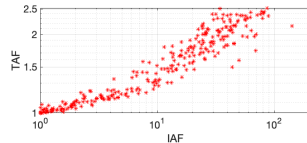
Monte Carlo: the background decides, not the seed



IAF vs. C



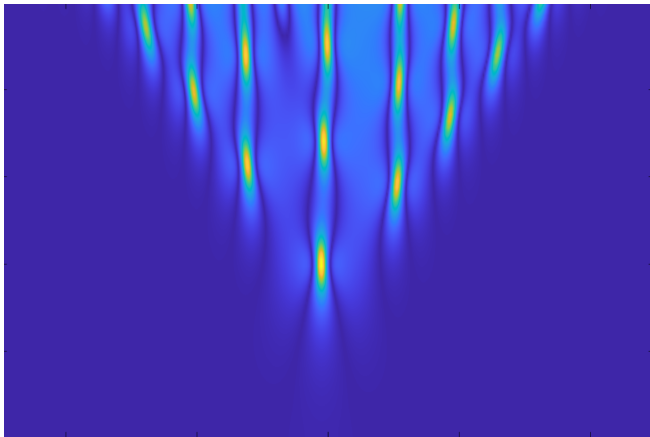
TAF vs. C



TAF vs. IAF

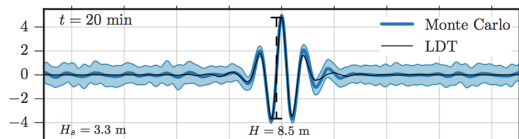
- 310 randomised initial conditions, $C \sim \mathcal{U}[0.9, 1.9]$.
- Amplification factors depend essentially **only on the background Γ** (i.e. on C), *not* on the initial inhomogeneity.
- $\text{TAF}(C)$ grows *quadratically*: **no jump** at onset $C \approx 1$; a substantial effect ($\text{TAF} \approx 1.4$, i.e. +40% variance) only from $C \approx 1.35$ or larger.
- **Physically plausible values:** $\text{TAF} \approx 1.5$, much smaller than 2.84.

Ramifications for rogue waves

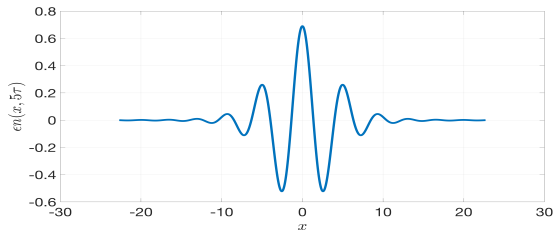


- Lattice of **hotspots of increased variance of the wavefield** found.
- The variance is not strong enough to guarantee Oceanic Rogue Waves, but it is strong enough to make them orders of magnitude more likely.
- Each peak is a “rogue wave” in a new nonlinear dispersive equation.
- Novel feature: it depends on Γ .
- Open questions: can we find an expression for it? Can we predict the structure of the lattice?

Conclusions



“Universal profile of rogue waves”
[van den Eijden et al. PNAS, 2018]



Profile predicted by the growth of Penrose-unstable modes
[A. et al., JOEME, 2017]

- The Alber equation presents a bifurcation that is only now being understood.
- It has clear applicability to oceanic rogue waves.
- Moreover, through generalized modulation instability, it generates new solutions with sharply localized features.

Thank you!

Appendix: initialization of the auxiliary variable

The first step needs $\Phi_{i,j}^{-\frac{1}{2}}$. Two choices:

$$\textbf{Naive: } \Phi_{i,j}^{-\frac{1}{2}} = U_{i,i}^0 - U_{j,j}^0.$$

$$\textbf{Advanced: } \begin{cases} \Phi_{i,j}^* = U_{i,i}^0 - U_{j,j}^0, \\ \left(I + i\frac{p\tau}{4}\mathfrak{D}_H + i\frac{q\tau}{4}\mathfrak{F}^* \right) \text{vec}(U_{i,j}^{-\frac{1}{4}}) = \text{vec}(U_{i,j}^0 - i\frac{q\tau}{4}\Gamma(x_i - y_j)\Phi_{i,j}^*), \\ U_{i,j}^{-\frac{1}{2}} = 2U_{i,j}^{-\frac{1}{4}} - U_{i,j}^0, \quad \Phi_{i,j}^{-\frac{1}{2}} = U_{i,i}^{-\frac{1}{2}} - U_{j,j}^{-\frac{1}{2}}. \end{cases}$$

- **Naive** (Besse): 2^{nd} order in u , but only 1^{st} **order in time** for ϕ .
- **Advanced** (half-step predictor): full order in *both* u and ϕ .
- The main variable u is largely insensitive to the error in ϕ (an averaging effect).

Appendix: experimental order of convergence

Time refinement ($h = 0.04$ fixed)

naive ($\text{EOC}_\phi \approx 1$):

τ	$\mathcal{E}_u$	EOC_u	$\mathcal{E}_\phi$	EOC_ϕ
3.0e-2	5.3e-3	—	9.6e-2	—
2.1e-2	2.7e-3	1.96	6.6e-2	1.06
1.5e-2	1.3e-3	2.05	4.6e-2	1.04
1.1e-2	6.7e-4	1.99	3.2e-2	1.03

advanced:

τ	$\mathcal{E}_u$	EOC_u	$\mathcal{E}_\phi$	EOC_ϕ
3.0e-2	4.6e-3	—	1.1e-2	—
2.1e-2	2.3e-3	1.96	5.4e-3	2.01
1.5e-2	1.1e-3	2.06	2.7e-3	1.99
1.1e-2	5.7e-4	1.98	1.3e-3	2.01

Space refinement ($\tau = 5 \cdot 10^{-4}$ fixed)

naive:

h	$\mathcal{E}_u$	EOC_u	$\mathcal{E}_\phi$	EOC_ϕ
0.40	1.3e-2	—	2.4e-2	—
0.34	6.7e-3	3.89	1.2e-2	3.78
0.28	3.3e-3	4.07	6.3e-3	3.88
0.24	1.7e-3	4.04	3.4e-3	3.55

advanced:

h	$\mathcal{E}_u$	EOC_u	$\mathcal{E}_\phi$	EOC_ϕ
0.40	1.3e-2	—	2.3e-2	—
0.34	6.7e-3	3.89	1.2e-2	3.86
0.28	3.3e-3	4.07	5.9e-3	4.06
0.24	1.7e-3	4.04	2.9e-3	4.03

Appendix: the a priori $H^1 \mathfrak{S}^1$ bound (no smallness)

Kinetic energy $\mathcal{K}(t) := \text{tr}(-\Delta \gamma(t))$; conserved energy and mass

$$E(\gamma) = p \text{tr}(\Delta \gamma) + \frac{q}{2} \|\rho_\gamma\|_{L^2}^2, \quad \mathcal{M} := \|\gamma\|_{\mathfrak{S}^1} = \|\rho_\gamma\|_{L^1} \quad (\gamma \geq 0).$$

Focusing ($pq > 0$). Energy conservation $\Rightarrow |p| \mathcal{K}(t) \leq |E(\gamma_0)| + \frac{|q|}{2} \|\rho_\gamma\|_{L^2}^2$, so it is enough to bound $\|\rho_\gamma\|_{L^2}^2$ by $\mathcal{K}$. Two inequalities (the first *needs* $\gamma \geq 0$):

$$\underbrace{\int_{\mathbb{T}} |\nabla \sqrt{\rho_\gamma}|^2 \leq \text{tr}(-\Delta \gamma),}_{\text{Hoffmann-Ostenhof}} \quad \underbrace{\|f\|_{L^4}^4 \leq \frac{1}{|\mathbb{T}|} \|f\|_{L^2}^4 + 2 \|f\|_{L^2}^3 \|\nabla f\|_{L^2}}_{\text{Gagliardo-Nirenberg on } \mathbb{T}}.$$

Take $f = \sqrt{\rho_\gamma}$, use $\|\sqrt{\rho_\gamma}\|_{L^2}^2 = \|\rho_\gamma\|_{L^1} = \mathcal{M}$:

$$\|\rho_\gamma\|_{L^2}^2 = \|\sqrt{\rho_\gamma}\|_{L^4}^4 \leq C \mathcal{M}^{3/2} \sqrt{\mathcal{K}(t)} + C \frac{\mathcal{M}^2}{|\mathbb{T}|} \quad (\text{sublinear in } \mathcal{K}).$$

Feed back into the energy inequality:

$$\mathcal{K}(t) \leq A + B \sqrt{\mathcal{K}(t)} \implies \mathcal{K}(t) \leq \bar{Y} \quad \forall t \quad (\text{no smallness}).$$

Defocusing ($pq < 0$). Kinetic and potential terms share a sign, so $|p| \mathcal{K}(t) \leq |E(\gamma_0)|$ *directly*.

Towards nonlinear stability

Take a homogeneous background $\Gamma = \Gamma^* \geq 0$ that is **Penrose-stable** (linearly stable: no unstable modes, with a stability margin $\kappa > 0$). Perturb it by a small, self-adjoint U_0 with $\|U_0\|_{H^1 \mathfrak{S}^1} \leq \varepsilon$ and $\gamma_0 = \Gamma + U_0 \geq 0$.

Theorem (Polynomial-in- t control near Penrose-stable Γ [A. arXiv:2604.16998])

The solution $\gamma(t)$ exists globally in $C(\mathbb{R}; H^1 \mathfrak{S}^1(\mathbb{T}))$ and

$$\|\gamma(t) - \Gamma\|_{H^1 \mathfrak{S}^1} \leq 2 C_\star(\Gamma) \langle t \rangle^2 \varepsilon \quad \text{for all } |t| \leq T_\star, \quad T_\star = c_\Gamma \varepsilon^{-1/5},$$

with $C_\star(\Gamma), c_\Gamma > 0$ independent of ε . In particular $\|\gamma(t) - \Gamma\|_{H^1 \mathfrak{S}^1} \lesssim \varepsilon^{3/5}$ on that window.

- Bare well-posedness gives only the exponential $\|\gamma(t) - \Gamma\|_{H^1 \mathfrak{S}^1} \leq \varepsilon e^{Ct}$; Penrose stability upgrades this to **polynomial** growth on the long window $|t| \lesssim \varepsilon^{-1/5}$.
- On Penrose-*unstable* backgrounds the exponential is attained ($\|\gamma(t) - \Gamma\| \gtrsim \varepsilon e^{ct}$) — this is the classical modulation instability.
- Falls short of genuine orbital stability ($\|\gamma(t) - \Gamma\| = o(1)$ as $t \rightarrow \infty$), which we conjecture but do not prove.